## SEMESTER LEARNING ACTIVITY PLANS (SLAP) SEMESTER ODD 2022/2023



Physics Undergraduate Study Program Physics Department Introduction to Astrophysics and Cosmology MFF 4043/ 3 Credits

Lecturer Coordinator:

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UNIVERSITAS GADJAH MADA FACULTY OF MATHEMATICS AND NATURAL SCIENCE 2022

## Universitas Gadjah Mada

Faculty of Mathematics and Natural Science Physics Department / Physics Undergraduate Study Program Semester ODD 2022/2023 **Document Number :** 

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## SEMESTER LEARNING ACTIVITY PLANS (SLAP)

| Code                                  | Course<br>Name   | Credits (Credits)  |  | Semester   | Status      | Pr                     | erequisite          |
|---------------------------------------|--|--|--|--|-------------|------------------------|---------------------|
| MFF 4043                              | Introductio  | <i>T: 3</i>  | <i>P</i> :                                     | ODD  | Elective    | Core and I             | Particle oh Physics |
|                                       | n to   |  |  |  |             | <b>I</b> (1            | MFF 2205)           |
|                                       | Astrophysic  |  |  |  |             |                        |                     |
|                                       | s and  |  |  |  |             |                        |                     |
|                                       | Cosmology  |  |  |  |             |                        |                     |
| Short<br>Description                  | The Introductory Course in Astrophysics and Cosmology is a 3 Credits elective course in the 2021 curriculum of the Gadjah Mada University Physics Study Program, which can be taken in odd semesters. To take this course, students are recommended to complete the Theory of Relativity and Core Physics and particle I courses. In the 2021 Curriculum of the Physics Undergraduate Study Program, this course is associated with competencies in Aspects of Knowledge (PLO 2), Aspects of Special Skills (PLO 4), and Aspects of Long-Life Learning/self-development (PLO 5). |  |  |  |             |                        |                     |
| Program                               | PLO 2  | Knowledge. Able to explain theoretical concepts and principles of classical and modern physics and able to apply basic concepts of physics and related mathematical methods in finding solutions to physical problems. |  |  |             |                        |                     |
| Learning<br>Outcomes<br>(PLO) Imposed | PLO 4  | Special Skills. Able to design and carry out experiments/theoretical reviews, able to identify a physical problem based on the results of observations and experiments, and able to operate related technologies.      |  |  |             |                        |                     |
| on the Course                         | PLO 5  | Long Life Learning. Able to analyze various alternative solutions to physical problems and conclude them for appropriate decision-making, both in familiar and new problems.   |  |  |             |                        |                     |
|                                       | After comple   | eting this course, students are expected to be able to:  |  |  |             |                        |                     |
| Course                                | <i>CO1</i>   | Knowing and U  | Jnder  | standing Stellar As  | strophysics |                        |                     |
| Outcomes (CO)                         | <i>CO2</i>   | Knowing and understanding the Galaxy   |  |  |             |                        |                     |
|                                       | СОЗ  | Knowing and understanding Cosmology  |  |  |             |                        |                     |
|                                       |  | Learning Materials Learning Methods Time A   |  |  |             | <b>Time Allocation</b> |                     |
| The Correlation<br>of CO to           | CO 1   | Stellar Astroph<br>radiation from<br>brightness, star<br>radiation, stella<br>and absolute m   | ysics:<br>the sta<br>color<br>r dist<br>agnitu | Continuous<br>ars (star<br>; black body<br>ance, luminosity,<br>ude) |             |                        | 3X50 minutes        |
| Learning                              | CO 1   | Stellar Astroph  | ysics:   | b. Line  |             |                        |                     |
| Materials and                         |  | spectrum and it  | s for  | nation (Bohr-  |             |                        |                     |
| Methods, and                          |  | Summerfield at   | omic   | model, line  |             |                        |                     |
| Time Allocation                       |  | spectrum forma   | ation,   | Hertzsprung-   |             |                        | 3X50 minutes        |
|                                       |  | Russel diagram   | ), Tel   | lescopes and   |             |                        | 52150 minutes       |
|                                       |  | other detectors  | (optio   | cal telescopes,  |             |                        |                     |
|                                       |  | other waveleng<br>and gravitation  | th reg<br>al wa                                | gions, neutrinos,<br>ves   |             |                        |                     |

| CO 1        | Stellar Astrophysics: d. Binary stars  |                              |               |
|-------------|--|------------------------------|---------------|
|             | and star parameters (Kepler's Laws,  |                              | 3X50 minutes  |
|             | mass-luminosity relation, star radius)   |                              |               |
| <i>CO</i> 1 | Stellar Astrophysics: e. Star  |                              |               |
|             | atmosphere and radiation transport,  |                              |               |
|             | main sequence stars and stellar  |                              |               |
|             | structure (stellar structure equations,  |                              | 3X50 minutes  |
|             | Eddington luminosity and convective  |                              | 52150 minutes |
|             | instability, Eddington standard  |                              |               |
|             | model, stellar stability, stellar  |                              |               |
|             | variables).  |                              |               |
| CO 1        | Stellar Astrophysics: Nuclear  |                              |               |
|             | Processes in stars (Sources of energy  |                              |               |
|             | in stars, fundamental interactions,  |                              | 3X50 minutes  |
|             | thermonuclear reactions, significant   |                              | 51100 minutes |
|             | nuclear combustion reactions, solar  |                              |               |
|             | neutrinos)   |                              |               |
| <i>CO</i> 1 | Stellar Astrophysics: Endpoints of   |                              |               |
|             | stellar evolution (Sirius B  |                              |               |
|             | observations, degenerative Fermi gas   |                              | 3X50 minutes  |
|             | pressure, Chandrasekar white dwarfs  |                              |               |
|             | and limits, supernovae, pulsars, and   |                              |               |
| <u> </u>    | Neutron stars)   |                              |               |
| <i>C01</i>  | Stellar Astrophysics: Black Holes  |                              |               |
|             | (Schwarzschild metric, gravitational   |                              | 2V50          |
|             | radiation from pulsars,  |                              | 3X30 minutes  |
|             | Disely Holes)  |                              |               |
|             | Midtorm over/Dreject Teck D  | agulta/Casa Analygia Dagulta |               |
|             | Galaxiaa: Star formation and   | esuits/Case Analysis Results |               |
| <i>CO</i> 2 | intersteller medium (intersteller dust   |                              | 3V50 minutos  |
|             | interstellar medium (interstellar dust,  |                              | SASO minutes  |
| <i>CO</i> 2 | Galaxias: Star clusters (avolved   |                              |               |
|             | dobular clusters, virial masses  |                              |               |
|             | Hetzsprung-Russel masses in  |                              | 3X50 minutes  |
|             | clusters)  |                              |               |
| <u>CO 2</u> | Galaxy: Galaxy (Milky Way, regular   |                              |               |
| 002         | and active galaxies non-thermal  |                              | 3X50 minutes  |
|             | radiation)   |                              | 57150 minutes |
| CO 3        | Cosmology: a An introduction to the  |                              |               |
| 000         | universe on a large scale (problems  |                              |               |
|             | on the static Newtonian universe.  |                              | 3X50 minutes  |
|             | cosmological principles, expansion   |                              |               |
|             | of the universe and Hubble's laws).  |                              |               |
| CO 3        | Cosmology: b. Cosmological model   |                              |               |
|             | (Friedmann's equations, scale  |                              |               |
|             | dependencies on various forms of   |                              | 3X50 minutes  |
|             |  |                              |               |
|             | energy, cosmological model with one  |                              |               |
|             | energy, cosmological model with one<br>energy component, ACDM model)                                 |                              |               |
| CO 3        | energy, cosmological model with one<br>energy component, ACDM model)<br>Cosmology: c. Young universe |                              | <u></u>       |

|  | Bang Nucleosynthesis, structure   |                          |                         |              |              |            |
|--|---|--------------------------|-------------------------|--------------|--------------|------------|
|  | Final exams/ Project Task Results/Case Analysis Results   |                          |                         |              |              |            |
| Learning   | SCL (Student Centered Learning): Project-based learning (Team-based Project)/Case-based   |                          |                         |              |              |            |
| Methods  | learning/PBL/other SCL methods  |                          |                         |              |              |            |
| Student<br>Learning<br>Experience  | Listen, ask, answer questions and discuss   |                          |                         |              |              |            |
| Access to<br>Learning<br>Media/ LMS<br>and Offline and<br>Online<br>Percentage | Offline (LCD, PPT Slide, Whiteboard, Laptop) and Online (Zoom Meeting, Google Meet, Google Classroom)   |                          |                         |              |              |            |
|  | Assessment<br>Methods   | Assessment<br>Percentage | Criteria/<br>Indicators | CO1          | CO2          | CO3        |
| Assessment   | Participatory<br>Activity*  | Tercentage               |                         |              |              |            |
|  | Project<br>Results/ Case<br>Study Results/<br>PBL Results*  |                          |                         |              |              |            |
| Methods and<br>Synchronizatio  | Cognitive   | 10                       |                         |              | 1 1          | - <u> </u> |
| n with CO  | Assignment  | 10                       |                         | <u> </u>     | N            | <b>√</b>   |
|  | Quiz  | 10                       |                         | ٧            | N            | N          |
|  | Exam  | 40                       |                         | $\checkmark$ | $\checkmark$ |            |
|  | Final Exam  | 40                       |                         |              | √            | √          |
|  | Total   | 100                      |                         |              |              |            |
|  | <sup>(7)</sup> can also be obtained from the Midterm or Final Exam as the result of participatory activities or process study results. According to IKU 7, the percentage of project results/ case study/ PBL results least 50%.  |                          |                         |              |              |            |
| References   | <ol> <li>Main References;         <ol> <li>Arnab Rai Choudhuri, Astrophysics for Physicists, 2010, Cambridge University Press.</li> <li>Francis Leblanc, An Introduction to Stellar Astrophysics, 2010, John Wiley and Sons, Ltd.</li> <li>David Lyth, Cosmology For Physicists, 2017, Taylor &amp; Francis Group, LLC.</li> <li>Ryden, B. Introduction of Cosmology, 2016, Department of Astronomy, The Ohio State University.</li> </ol> </li> <li>Raine, D.J &amp; Thomas, E.G, An Introduction To The Science Of Cosmology, 2001, IOP Publishing.</li> <li>M. Kachelrieß, A Concise Introduction to Astrophysics, 2011, Institutt for fysikk NTNU, Trondheim Norway.</li> </ol> |                          |                         |              |              |            |
| Lecturers<br>(Team<br>Teaching)  | <ol> <li>Dr. Eng. Rinto Anugraha NQZ, S.Si., M.Si.</li> <li>Romy Hanang Setya Budhi, S.Si., M,Sc., Ph.D.</li> </ol>   |                          |                         |              |              |            |

| Authorization | Date of<br>Drafting | Lecturer Coordinator                         | Head of<br>Curriculum<br>Committee | Head of Study Program                        |
|---------------|---------------------|--|------------------------------------|--|
|               |                     | Dr. Eng. Rinto Anugraha NQZ, S.Si.,<br>M.Si. |                                    | Dr. Eng. Ahmad<br>Kusumaatmaja, S.Si., M.Sc. |